

# Physics 566 - Quantum Optics

## Problem Set #2 - Solutions

### Problem 1

Two level atom. In RWA (in rotating frame)

$$\hat{H}_{\text{eff}} = -\frac{\hbar\Delta}{2} \hat{\sigma}_z - \frac{\hbar R}{2} \hat{\sigma}_x = -\frac{\hbar}{2} \vec{\Omega} \cdot \hat{\sigma}$$

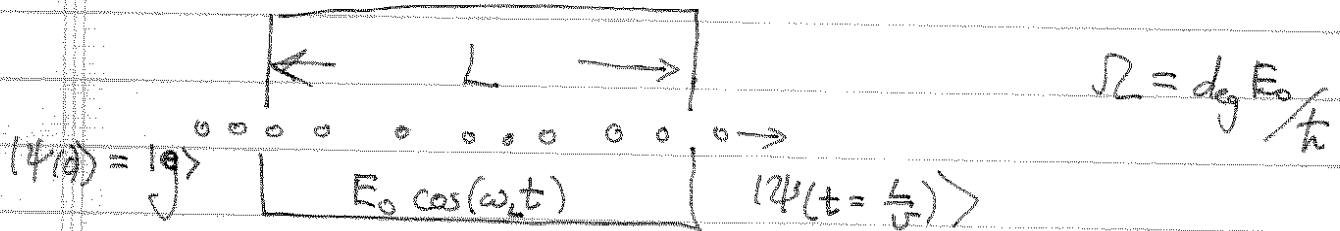
$$\vec{\Omega} = \Delta \vec{e}_z + R \vec{e}_x = \tilde{\Omega} \vec{e}_n$$

$$\tilde{\Omega} = \sqrt{R^2 + \Delta^2} \quad \vec{e}_n = \cos\theta \vec{e}_z + \sin\theta \vec{e}_x \quad \begin{aligned} \cos\theta &= \frac{\Delta}{\tilde{\Omega}} \\ \sin\theta &= \frac{R}{\tilde{\Omega}} \end{aligned}$$

Unitary evolution (in rotating frame)

$$\begin{aligned} \hat{U}(t) &= e^{-i\hat{H}_{\text{eff}}t/\hbar} = e^{i\frac{\tilde{\Omega}t}{2} \hat{\sigma}_n} = \cos\left(\frac{\tilde{\Omega}t}{2}\right) \hat{1} + i\sin\left(\frac{\tilde{\Omega}t}{2}\right) \hat{\sigma}_n \\ &= \cos\left(\frac{\tilde{\Omega}t}{2}\right) \hat{1} + i\sin\left(\frac{\tilde{\Omega}t}{2}\right) (\cos\theta \hat{\sigma}_z + \sin\theta \hat{\sigma}_x) \end{aligned}$$

(i) Rabi geometry



(a) Mono-energetic beam with velocity  $v$ , length chosen such that  $\Omega \frac{L}{v} = \pi$  (i.e.  $\pi$ -pulse for atoms on resonance)

Now we have the Rabi solution:

$$\begin{aligned}
 |\psi(t)\rangle &= \hat{U}(t) |g\rangle = \\
 &= \left( \cos\left(\frac{\tilde{\Omega}t}{2}\right) - i \cos\theta \sin\left(\frac{\tilde{\Omega}t}{2}\right) \right) |g\rangle + i \sin\theta \sin\frac{\tilde{\Omega}t}{2} |e\rangle
 \end{aligned}$$

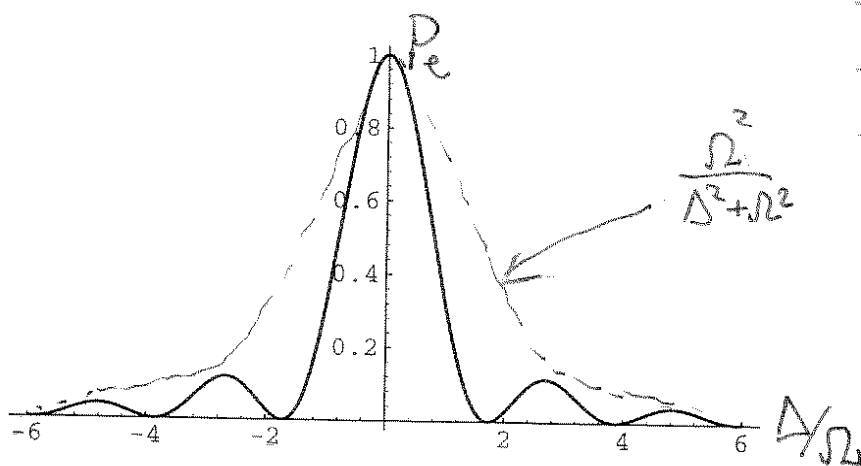
Thus the probability to be in the excited state after a time  $t = \frac{L}{v}$  is

$$P_e = \sin^2\theta \sin^2\left(\frac{\tilde{\Omega}L}{2v}\right) = \frac{\Omega^2}{\Delta^2 + \Omega^2} \sin^2\left(\frac{\sqrt{\Omega^2 + \Delta^2} L}{2v}\right)$$

$$\Rightarrow P_e = \frac{1}{1 + \left(\frac{\Delta}{\Omega}\right)^2} \sin^2\left(\left(1 + \frac{\Delta^2}{\Omega^2}\right)^{1/2} \frac{\pi}{4}\right)$$

(where I used  $\Omega L/v = \pi$ )

Plot of  $P_e$  as a function of  $\Delta = \omega_L - \omega_{eg}$  in units of  $\Omega$



Linewidth  $\Delta\omega \sim \Omega_L = \frac{\pi}{T}$  where  $T = \frac{L}{v}$

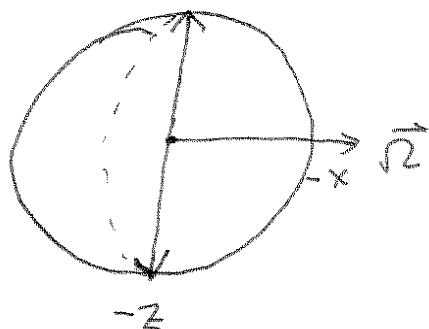
(d) Continued

We see that the linewidth is on the order

$$\Delta\omega \approx \Omega = \frac{\pi}{T} \quad \text{where } T = \frac{L}{v} = \text{interaction time}$$

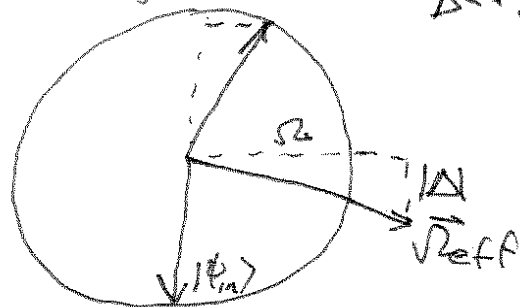
On the Bloch-sphere

Resonance  $\Delta=0$



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Slightly off-resonance  $\Delta \ll \Omega$



$$\vec{\Omega}_{eff} = \Omega \vec{e}_x + \Delta \vec{e}_z$$

From simple geometry  
we see that  $P_e = \left| \frac{\Omega_x}{\sqrt{\Delta^2 + \Omega^2}} \right|^2$   
(for  $\vec{\Omega}T = \pi$ )

It is clear from these sketches that when  $\Delta$  is on the order of  $\Omega$  population transfer to the excited state decreases substantially. Since we have fixed  $\Omega T = \pi$  we get the most sensitivity to  $\Delta$  by making  $T$  longer and thus  $\Omega$  smaller.

(b) Now suppose the atoms have a distribution of velocities characteristic of thermal beams

$$f(v) = 2 \frac{v^3}{v_0^4} e^{-\frac{v^2}{v_0^2}} \quad \text{where} \quad v_0 = \sqrt{\frac{2kT}{m}}$$

Note: for this distribution

- Most probable velocity:  $v_p = 1.22 v_0$
- Average velocity:  $\bar{v} = 1.33 v_0$
- rms velocity:  $\sqrt{\Delta v^2} = 1.42 v_0$

We have for a fixed velocity and detuning

$$P_e(A, v) = \frac{1}{1 + \left(\frac{A}{v_0}\right)^2} \sin^2\left(\left(1 + \frac{A^2}{v_0^2}\right)^{1/2} \frac{\Omega L}{2v}\right)$$

At zero detuning (on resonance):  $P_e(0, v) = \sin^2\left(\frac{\Omega L}{2v}\right)$

Now we must average over the velocity distribution:

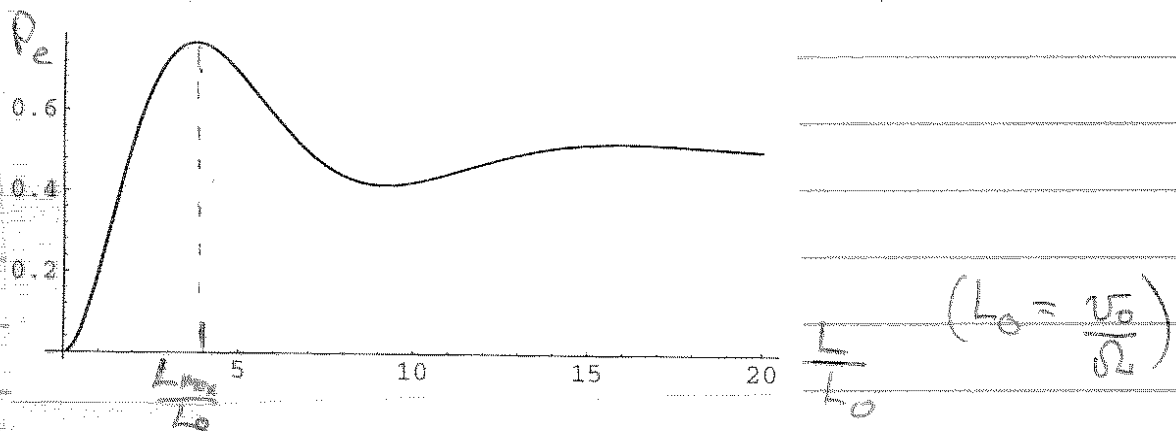
$$P_e = \int_0^{\infty} dv f(v) P_e(0, v) = \int_0^{\infty} dx f(x) \sin^2\left(\frac{\Theta_0}{2x}\right)$$

where  $x \equiv \frac{v}{v_0}$  (dimensionless velocity)

and  $\Theta_0 \equiv \frac{\Omega L}{v_0}$  (Angle turned by the Bloch vector for atoms traveling at  $v_0$  and  $\omega_1$  on resonance)

Note: Always re-express in terms of dimensionless variables

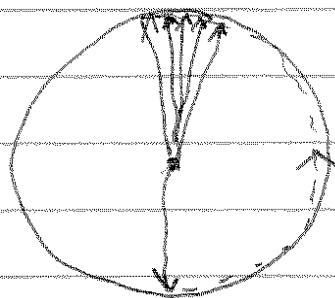
Plot (calculated with Mathematica)



The maximum  $P_e$  occurs at  $\Theta_0 = \frac{\Omega L_{\max}}{v_0} = 3.77$   
 or  $\frac{\Omega L_{\max}}{v_0} \approx 1.20 \pi$ . We can understand this

from the fact that the most probable speed is  
 $v_p \approx 1.2 v_0 \Rightarrow \frac{\Omega L_{\max}}{v_p} \approx \pi$

~~What~~ On the Bloch-sphere:



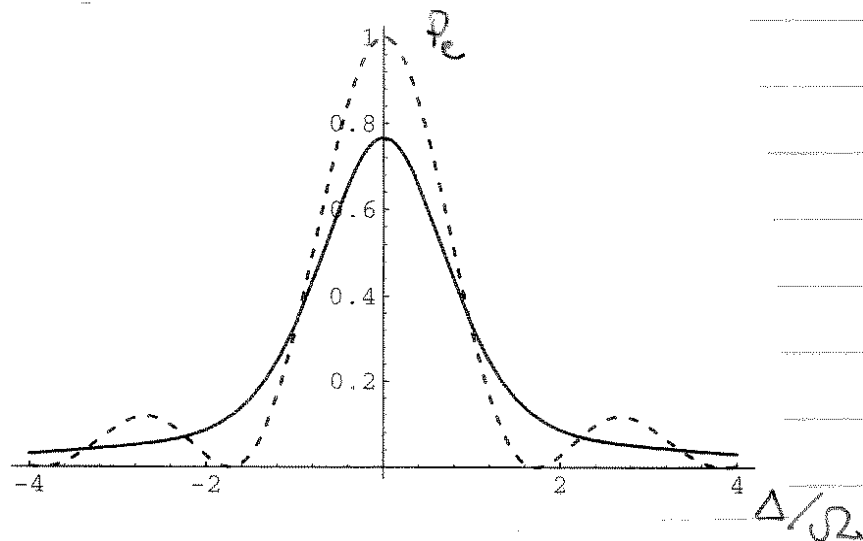
Distribution of velocities  
 $\Rightarrow$  Distribution of interaction times  
 $\Rightarrow$  Distribution of rotation angles

This is an example of "inhomogeneous" broadening. The "damped" Rabi oscillations plotted above are not due to dissipation. Rather they are due to the spread in rotation angle due to an inhomogeneity in the sample (here velocity).

Now set  $\theta_0 = 1.2\pi \approx 3.8$ , and  $\Delta \neq 0$

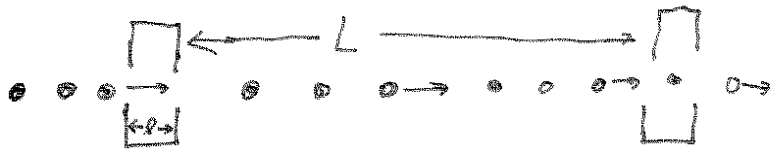
$$\Rightarrow P_e = \frac{1}{1 + \left(\frac{\Delta}{\Omega}\right)^2} \int_0^{\infty} \sin^2 \left( \left(1 + \frac{\Delta^2}{\Omega^2}\right)^{1/2} \frac{\theta_0}{2} \frac{1}{x} \right) (2x^3 e^{-x^2}) dx$$

Plot as a function of  $\Delta/\Omega$  with  $L = L_{\max}$   
(solid curve)



For reference I have added the curve for mono-energetic atom (dashed curve). We see that the spread in velocities leads to a broadening of the resonance lineshape.

(ii) Ramsey separated zone method



(c) Given  $|\psi(0)\rangle = |g\rangle$  and  $|\Delta| \ll \Omega$

In the interaction regions:  $\tilde{\Omega} \approx \Omega$      $\Theta \approx \pi/2$

$$\hat{U}_{int}(t) \approx \cos\left(\frac{\Omega t}{2}\right) \hat{I} + i \sin\left(\frac{\Omega t}{2}\right) \hat{\sigma}_x = e^{i\frac{\Omega t}{2} \hat{\sigma}_x}$$

In the free zone:  $\tilde{\Omega} = \Delta$      $\Theta = 0$

$$\hat{U}_{free}(t) = \cos\left(\frac{\Delta t}{2}\right) \hat{I} + i \sin\frac{\Delta t}{2} \hat{\sigma}_z = e^{i\frac{\Delta t}{2} \hat{\sigma}_z}$$

(free precession about  $-\hat{e}_z$  axis with frequency  $\Delta$  in the rotating frame)

$$\begin{aligned} |\psi(\tau = \frac{L}{v})\rangle &= \cos\left(\frac{\Omega L}{2}\right) |g\rangle + i \sin\left(\frac{\Omega L}{2}\right) |e\rangle \\ &= \frac{1}{\sqrt{2}} (|g\rangle + i |e\rangle) \end{aligned}$$

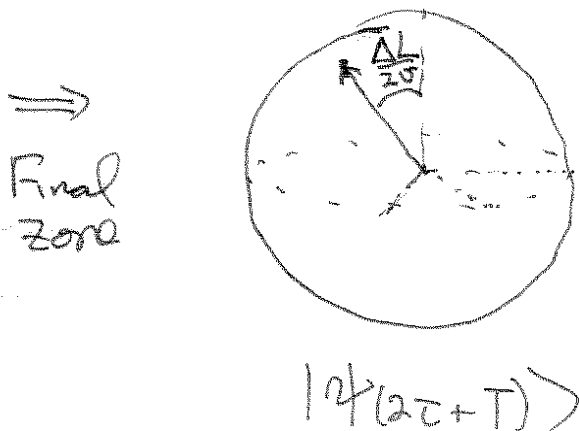
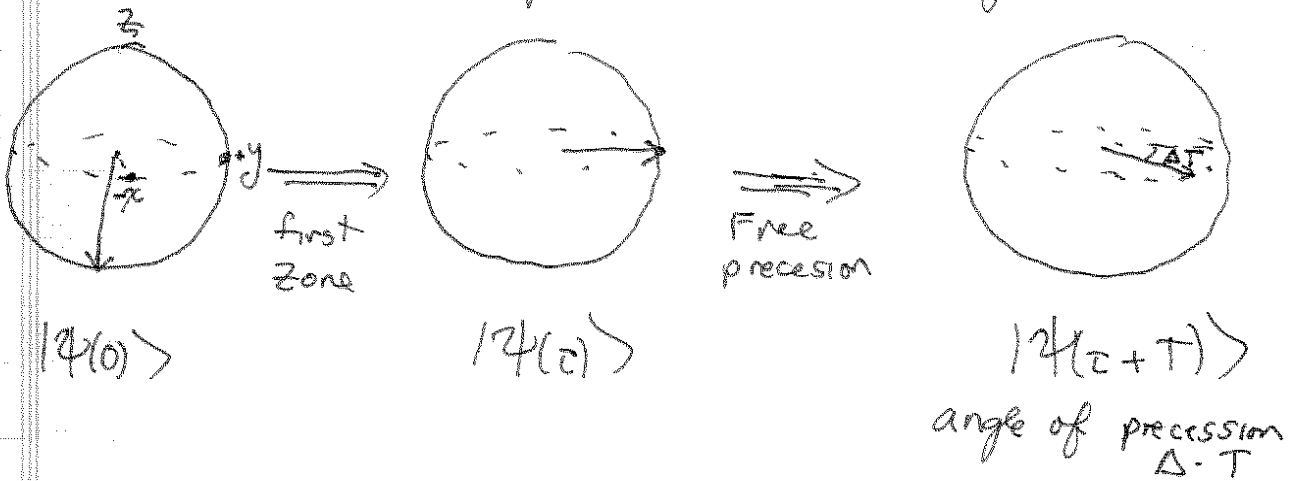
$$\begin{aligned} |\psi(\tau + T)\rangle &= \hat{U}_{free}(T) |\psi(\tau)\rangle = \frac{1}{\sqrt{2}} (e^{-i\frac{\Delta T}{2}} |g\rangle + i e^{i\frac{\Delta T}{2}} |e\rangle) \\ &= \frac{1}{\sqrt{2}} \left( e^{-i\frac{\Delta L}{2v}} |g\rangle + i e^{i\frac{\Delta L}{2v}} |e\rangle \right) \end{aligned}$$

(Next page)

$$\begin{aligned}
|\psi(2\tau+T)\rangle &= \hat{U}_{int}(\tau) |\psi(\tau+T)\rangle \\
&= \frac{1}{\sqrt{2}} \left( e^{-i\frac{\Delta L}{2\nu}} \hat{U}_{int}(\tau) |g\rangle + i e^{i\frac{\Delta L}{2\nu}} \hat{U}_{int}(\tau) |e\rangle \right) \\
&= \frac{1}{\sqrt{2}} \left( e^{-i\frac{\Delta L}{2\nu}} \frac{1}{\sqrt{2}} (|g\rangle + i|e\rangle) + i e^{i\frac{\Delta L}{2\nu}} \frac{1}{\sqrt{2}} (|e\rangle + i|g\rangle) \right) \\
&= \frac{1}{2} \left( e^{-i\frac{\Delta L}{2\nu}} - e^{i\frac{\Delta L}{2\nu}} \right) |g\rangle + \frac{i}{2} \left( e^{i\frac{\Delta L}{2\nu}} + e^{-i\frac{\Delta L}{2\nu}} \right) |e\rangle
\end{aligned}$$

$$\Rightarrow |\psi(2\tau+T)\rangle = -i \sin\left(\frac{\Delta L}{2\nu}\right) |g\rangle + i \cos\left(\frac{\Delta L}{2\nu}\right) |e\rangle$$

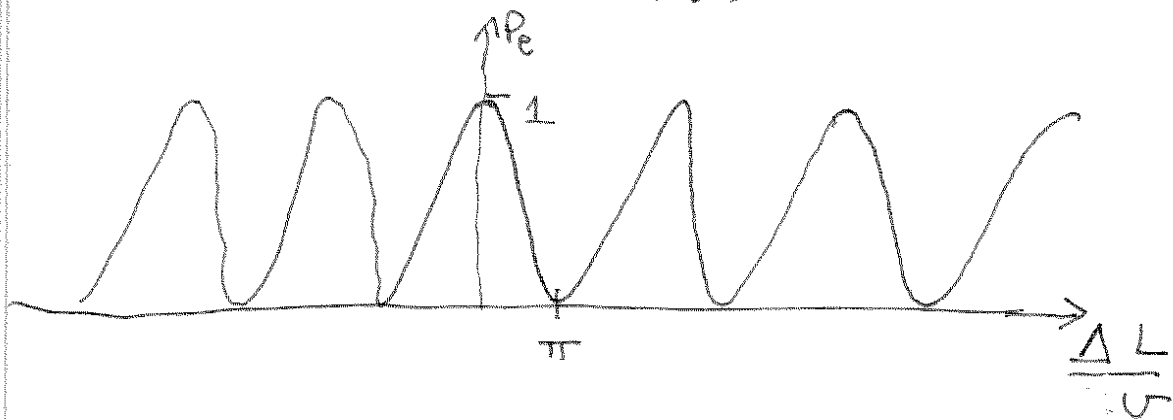
Block sphere in Rotating Frame



Final vector rotated towards north-pole. Final angle  $\theta = \pi - \frac{\Delta L}{2\nu}$

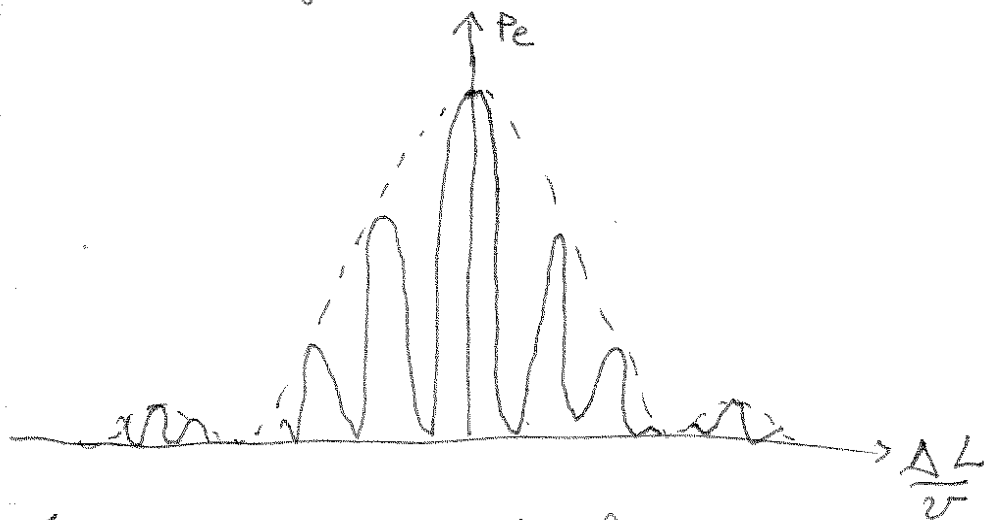


$$(d) P_e(t_{\text{final}}) = \cos^2\left(\frac{\Delta L}{2\nu}\right)$$



This is known as "Ramsey fringes"

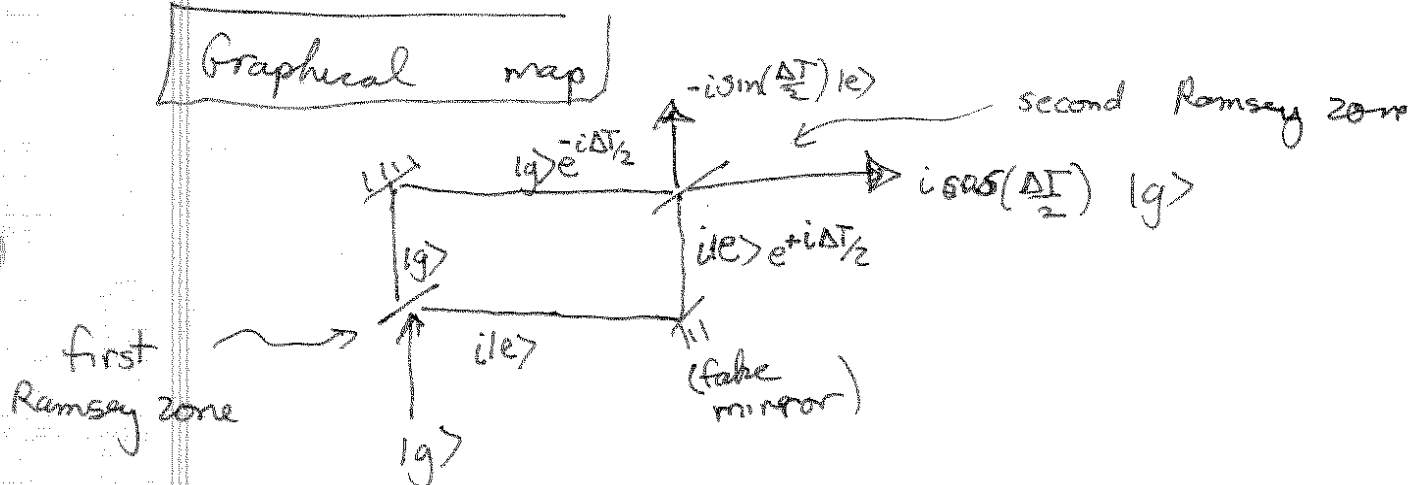
Note: We have assumed  $\Delta \ll \Omega$ . For the general solution we would have



The envelope of the fringe pattern is the Rabi pattern of part (a) due to the transit time broadening of width  $\Delta\omega \sim \frac{\nu}{l}$

The fringes themselves have a width on the order  $\Delta\omega \sim \frac{\nu}{L}$  which can be much smaller

(c) It is clear from the plots of part (d) that the Ramsey geometry results in a kind of interference pattern. We can think of the two "branches" of the wavefunction,  $|e\rangle$  or  $|g\rangle$ , as the two "arms" of an interferometer. The Ramsey zones produce  $\pi/2$  pulse are like "beam splitters". In the free interaction zone the two branches of the interferometer pick up a relative phase of  ~~$\Delta T$~~   $\Delta T$ .



The key to the Ramsey ~~method~~ method is to make  $T$  as big as possible. Then we can measure extremely small detunings  $\Delta$  because a measurable phase will accumulate (shown in the third mapping on the Bloch sphere). This is the basic physics behind the atomic clock. The idea is to lock a "local oscillator" to the natural oscillator of a two-level atom (Two hyperfine levels of cesium). When the frequency of the oscillator is on resonance ~~we~~ were there!

## Physics 56b

### Problem 2: Spin Precession in a magnetic field (10 points)

Consider a spin  $\frac{1}{2}$  particle such as an electron. It possesses an intrinsic magnetic moment described by the operator

$$\hat{\mu} = \gamma_s \hat{S}$$

where  $\gamma_s$  is "gyromagnetic ratio", and  $\hat{S}$  is the spin angular momentum operator. When placed in a magnetic field, the interaction Hamiltonian is given by

$$\hat{H} = -\hat{\mu} \cdot \vec{B} = -\gamma_s \vec{B} \cdot \hat{S} = -\gamma_s \sum_{j=1}^3 B_j \hat{S}_j$$

(a) Let us consider the dynamics in the ~~Schrodinger~~ Heisenberg picture. The equation of motion for the  $i$ th component of spin is:

$$\begin{aligned} \frac{d}{dt} \hat{S}_i &= \frac{1}{i\hbar} [\hat{S}_i, \hat{H}] = \frac{1}{i\hbar} [\hat{S}_i, -\gamma_s \sum_{j=1}^3 B_j \hat{S}_j] \\ &\stackrel{\text{Linearity}}{=} -\frac{\gamma_s}{i\hbar} \sum_{j=1}^3 B_j [\hat{S}_i, \hat{S}_j] = -\sum_{j,k=1}^3 \gamma_s B_j \hat{S}_k \\ &\qquad\qquad\qquad \sum_{k=1}^3 i\hbar \epsilon_{ijk} \hat{S}_k \qquad\qquad\qquad \epsilon_{ijk} \end{aligned}$$

Aside

Recall rule:  $(\vec{A} \times \vec{B})_i = \sum_{j,k=1}^3 \epsilon_{ijk} A_j B_k$

$$\Rightarrow \boxed{\frac{d}{dt} \hat{S} = -\vec{\Omega} \times \hat{S}} \quad \text{where} \quad \boxed{\vec{\Omega} = \gamma_s \vec{B}}$$

## Problem 2

(b) In spherical basis:

$$\begin{aligned}\hat{H} &= -\gamma_s \vec{B} \cdot \hat{\vec{S}} = -\frac{\hbar \gamma_s}{2} \vec{B} \cdot \hat{\vec{\sigma}} = -\frac{\hbar \Omega}{2} \cdot \hat{\vec{\sigma}} \\ &= -\frac{\hbar}{2} (\Omega_- \hat{\sigma}_+ + \Omega_+ \hat{\sigma}_- + \Omega_z \hat{\sigma}_z)\end{aligned}$$

$$\text{where } \Omega_{\pm} \equiv \Omega_x \pm i \Omega_y \quad \hat{\sigma}_{\pm} = \frac{\hat{\sigma}_x \pm i \hat{\sigma}_y}{2}$$

$$\begin{aligned}\text{Commutation relations } [\hat{\sigma}_z, \hat{\sigma}_{\pm}] &= \pm 2 \hat{\sigma}_{\pm} \\ [\hat{\sigma}_+, \hat{\sigma}_-] &= \hat{\sigma}_z\end{aligned}$$

Heisenberg equations of motion:

$$\boxed{\dot{\hat{\sigma}}_z = -\frac{i}{\hbar} [\hat{\sigma}_z, \hat{H}] = i(\Omega_- \hat{\sigma}_+ - \Omega_+ \hat{\sigma}_-)}$$

$$\dot{\hat{\sigma}}_+ = -\frac{i}{\hbar} [\hat{\sigma}_+, \hat{H}] = \frac{i}{2} ([\hat{\sigma}_+, \hat{\sigma}_-] \Omega_+ + [\hat{\sigma}_+, \hat{\sigma}_z] \Omega_z)$$

$$\boxed{\dot{\hat{\sigma}}_+ = \frac{i}{2} \Omega_+ \hat{\sigma}_z - i \Omega_z \hat{\sigma}_+}$$

Note: In the standard RWA Hamiltonian

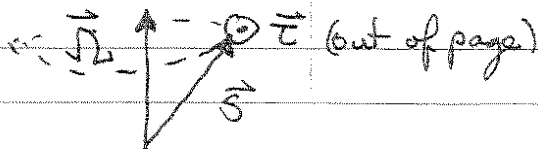
$$\begin{aligned}\Omega_+ = \Omega_- = \Omega & \quad \Omega_z = -\Delta \\ (\text{Rabi freq}) & \quad (\text{detuning})\end{aligned}$$

$$\Rightarrow \dot{\hat{\sigma}}_z = i\Omega (\hat{\sigma}_+ - \hat{\sigma}_-) \Rightarrow \langle \dot{\hat{\sigma}}_z \rangle = W(t)$$

$$\dot{\hat{\sigma}}_+ = i\frac{\Omega}{2} \hat{\sigma}_z + i\Delta \hat{\sigma}_+ \Rightarrow \langle \dot{\hat{\sigma}}_+ \rangle = U(t) + iV(t)$$

Physically: If  $\hat{S}$  were interpreted as a classical angular momentum vector then

$$\frac{d\vec{S}}{dt} = \vec{\tau} = \text{torque} = -\vec{\Omega} \times \vec{S}$$



Since  $\vec{\tau} \perp \vec{S}$ ,  $\vec{S}$  does not change length but instead, direction  $\Rightarrow \vec{S}$  precesses about  $-\vec{\Omega}$

Classically, this is what we expect for a classical magnetic moment in a static  $\vec{B}$  field. In that case the torque is  $\vec{\tau} = \vec{\mu} \times \vec{B} = \gamma \vec{L} \times \vec{B}$

$$\Rightarrow \frac{d\vec{L}}{dt} = \vec{\tau} = \vec{\mu} \times \vec{B} = \gamma \vec{L} \times \vec{B} = -\gamma \vec{B} \times \vec{L} = -\vec{\Omega} \times \vec{L}$$

The precession frequency  $\Omega = \gamma B$  is known as the Larmor frequency

(c) Solving for  $\hat{S}(t)$ : Take second derivative assuming static  $\vec{B}$  field

$$\frac{d^2 \hat{S}}{dt^2} = -\vec{\Omega} \times \frac{d\hat{S}}{dt} = \vec{\Omega} \times (\vec{\Omega} \times \frac{d\hat{S}}{dt}) = \vec{\Omega}(\vec{\Omega} \cdot \hat{S}) - \Omega^2 \hat{S}$$

For the case  $\vec{\Omega} = \Omega \vec{e}_z$  (magnetic field in z-direction)

$$\frac{d\hat{S}_x}{dt} = -\Omega \hat{S}_y \quad \frac{d\hat{S}_y}{dt} = \Omega \hat{S}_x \quad \frac{d\hat{S}_z}{dt} = 0$$

$$\frac{d^2 \hat{S}_x}{dt^2} = -\Omega^2 \hat{S}_x \quad \frac{d^2 \hat{S}_y}{dt^2} = -\Omega^2 \hat{S}_y$$

(Next Page)

Thus  $\hat{S}_x$  and  $\hat{S}_y$  satisfy the SHO equation and  $\hat{S}_z$  is unchanged

$$\Rightarrow \boxed{\hat{S}_z(t) = \hat{S}_z(0)}$$

$$\hat{S}_x(t) = \hat{S}_x(0) \cos \Omega t + \frac{1}{\Omega} \left. \frac{d\hat{S}_x}{dt} \right|_0 \sin \Omega t$$

$$\Rightarrow \boxed{\hat{S}_x(t) = \hat{S}_x(0) \cos \Omega t + \hat{S}_y(0) \sin \Omega t}$$

$$\hat{S}_y(t) = \hat{S}_y(0) \cos \Omega t + \frac{1}{\Omega} \left. \frac{d\hat{S}_y}{dt} \right|_0 \sin \Omega t$$

$$\Rightarrow \boxed{\hat{S}_y(t) = \hat{S}_y(0) \cos \Omega t - \hat{S}_x(0) \sin \Omega t}$$

Bloch vector  $\vec{Q}(t) = \langle \vec{\sigma}(t) \rangle$

in Heisenberg state  $| \downarrow \rangle = | -z \rangle$  (time independent)

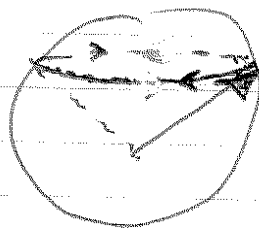
$$\Rightarrow Q_z(t) = \langle \sigma_z(t) \rangle = \langle -\frac{1}{2} \sigma_z(0) | -z \rangle = -1$$

$$Q_x(t) = \langle \sigma_x(t) \rangle = \cos \Omega t \langle \frac{1}{2} \sigma_x(0) | -z \rangle + \sin \Omega t \langle -\frac{1}{2} \sigma_y(0) | -z \rangle = 0$$

Similarly  $Q_y(t) = 0 \Rightarrow \downarrow$  Bloch sphere  
Eigenstate  $\Rightarrow$  No motion

If we started in an arbitrary state

$$| \uparrow_n \rangle = \cos \frac{\theta}{2} | \uparrow_z \rangle + e^{i\phi} \sin \frac{\theta}{2} | -z \rangle$$

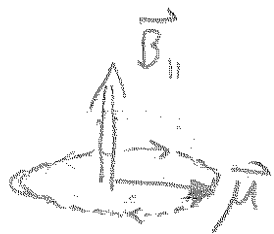


Bloch vector

Larmor precesses about  $-\vec{e}_z$  with frequency  $\Omega$

### Problem 3: Inhomogeneous broadening

(a) "Free induction" = free precession of the Bloch vector in  $x$ - $y$  plane



$$\Omega_n = \gamma |\vec{B}_{||}|$$

Larmor precession frequency

In the absence of dissipation, and for a fixed  $\vec{B}_{||}$ , given  $|\psi(0)\rangle = |+\rangle$

$$\langle \hat{\sigma}_x \rangle = \cos \Omega_n t = \cos(\gamma B_{||} t)$$

(i) Now we must average over the distribution of  $B_{||}$

$$\langle \hat{\sigma}_x \rangle = \int_{-\infty}^{\infty} dB_{||} \cos(\gamma B_{||} t) P(B_{||})$$

$$= \text{Re} \left[ \int_{-\infty}^{\infty} dB_{||} e^{-i(\gamma B_{||} t)} P(B_{||}) \right]$$

Fourier transform of  $P(B_{||})$

Gaussian distribution  $P(B_{||}) = \frac{1}{\sqrt{2\pi} \Delta B_{||}} e^{-\frac{(B_{||} - B_0)^2}{2\Delta B_{||}^2}}$

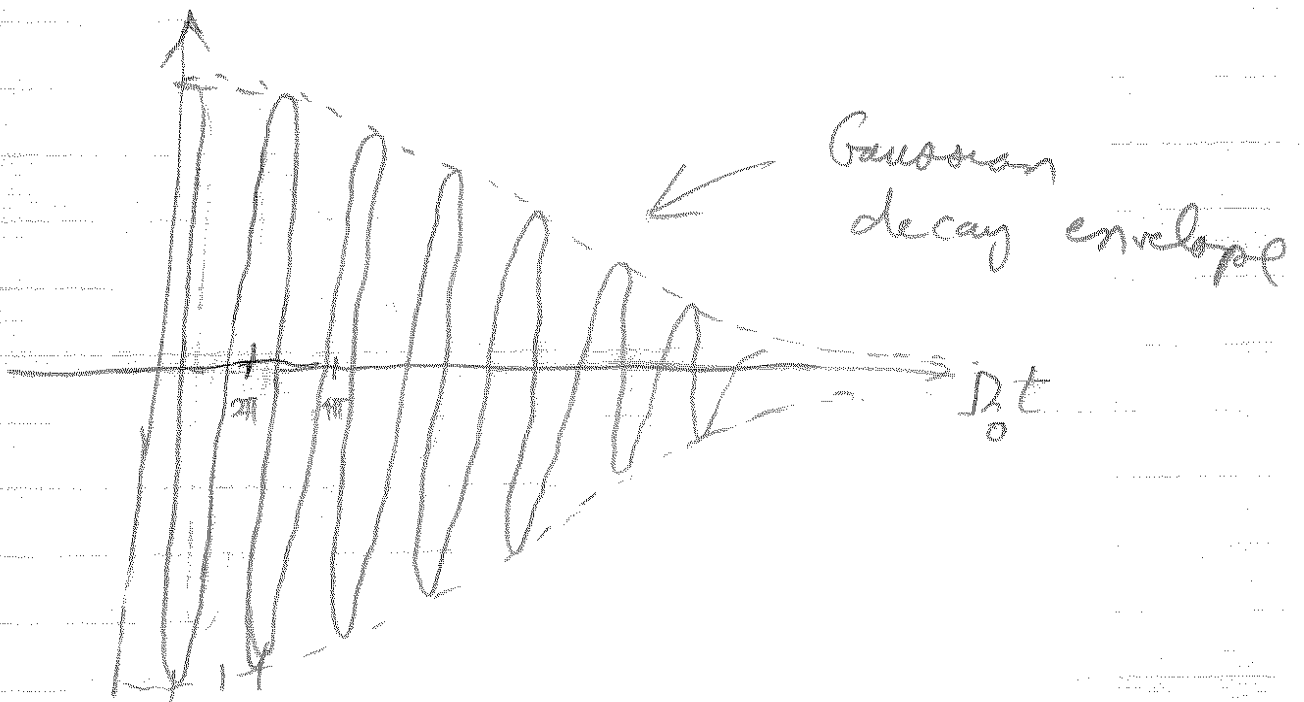
with  $\Delta B_{||} \ll B_0$

Change variables  $\Omega = \delta B_{11}$

$$\Rightarrow P(\Omega) = \frac{1}{\sqrt{2\pi} \Delta\Omega} e^{-\frac{(\Omega - \Omega_0)^2}{2\Delta\Omega^2}}$$

$$\int_{-\infty}^{\infty} d\Omega e^{-i\Omega t} P(\Omega) = e^{-i\Omega_0 t} \underbrace{\left( \int_{-\infty}^{\infty} d\bar{\Omega} e^{-i\bar{\Omega} t} \frac{e^{-\frac{\bar{\Omega}^2}{2\Delta\Omega^2}}}{\sqrt{2\pi} \Delta\Omega} \right)}_{\substack{\uparrow \\ \text{shift origin}}} e^{-\frac{(\Delta\Omega t)^2}{2} \frac{\sqrt{\Delta\Omega}}{\sqrt{2\pi}}}$$

$$\Rightarrow \langle \hat{\sigma}_x(t) \rangle = \frac{e^{-\frac{(\Delta\Omega t)^2}{2}}}{\sqrt{2\pi} \Delta\Omega} \cos(\delta B_0 t)$$



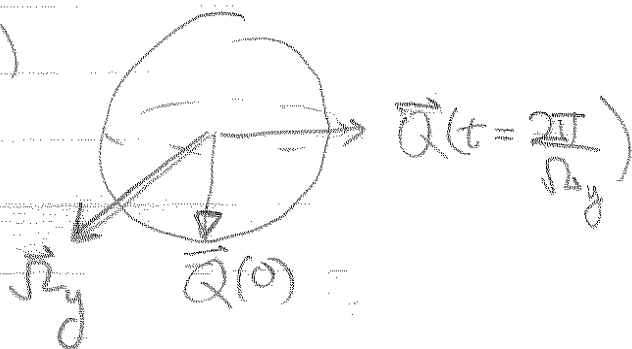


(ii) The decay here is due solely to the different spins getting out of phase.

The characteristic decay time is set by the width of the spectral distribution.

$$T_2^* \sim \frac{1}{\Delta\omega} = \frac{1}{\gamma\Delta B_{||}}$$

(iii)



By apply a pulse with  $\vec{\Omega}$  along  $\vec{e}_y$

such that  $\Omega_y T = \frac{\pi}{2}$

we rotate the spin to be along  $\vec{e}_x$ .

The problem is that  $\Omega_y$  is not well defined

because of the inhomogeneity. However, the effect of inhomogeneity is cause the different spins do get out of phase. So, if the time of the pulse  $T \ll T_2^*$  everything basically stays in phase. Thus we must have

$$\frac{\pi}{2\Omega_y} \ll T_2^* \Rightarrow \left[ \Omega_y \gg \frac{\pi}{2T_2^*} \right]$$

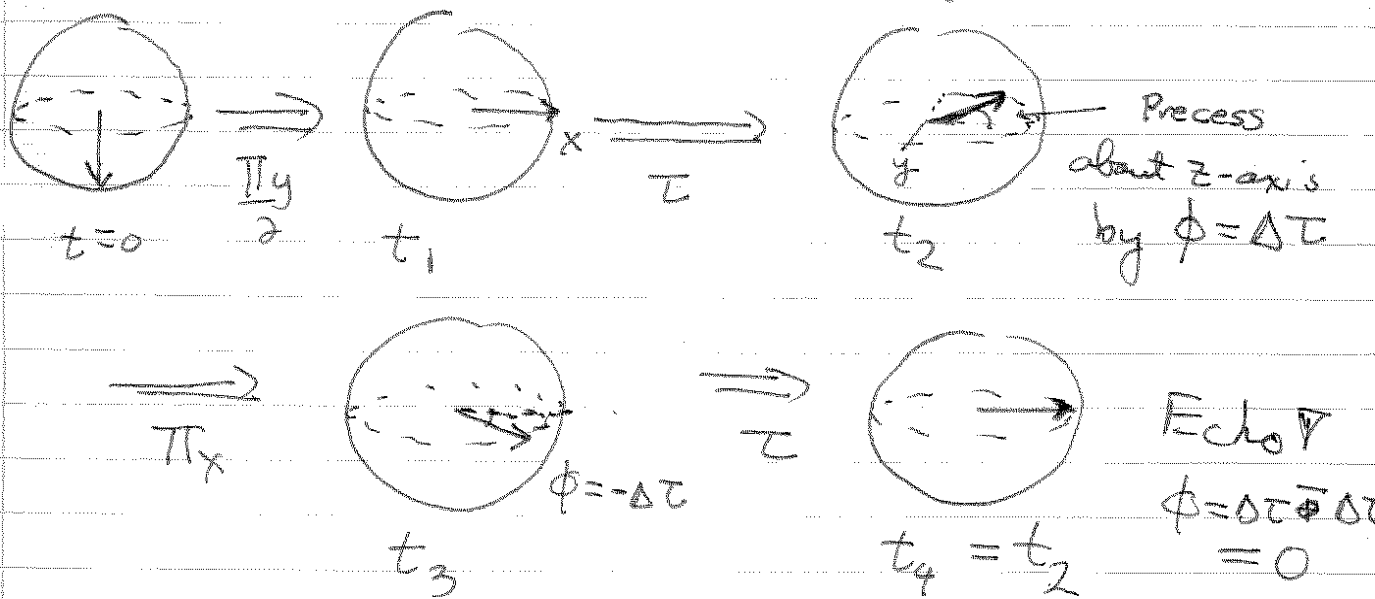
## (b) Spin echo

Consider the following pulse sequence



Let us consider a single spin with resonance frequency detuned  $\Delta$  from resonance of carrier. This pulse is assumed to be sufficiently broadband to bring the spin into  $x-y$  plane, independent of  $\Delta$ .

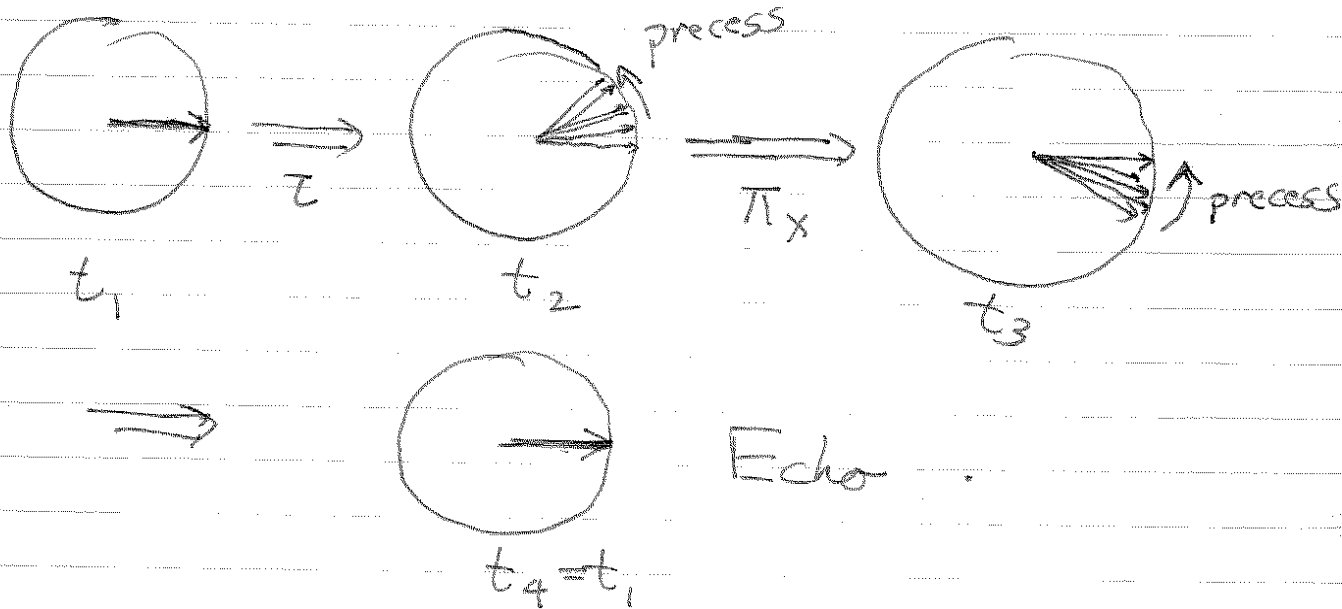
Sequence on Bloch-sphere (in rotating frame)



During the time intervals  $\tau$ , the spin precesses at frequency  $\Delta$  (in rotating frame) about  $z$ -axis. The  $\pi$  pulse effectively puts the spin "behind the starting line" by an amount exactly equal the amount it precessed. Thus after the second free precession period  $\tau$  the spin returns to the  $x$ -axis (Next page)

Let us now consider an ensemble of spins with a spread in  $\Delta$ ,  $\Delta\omega_{inh}$ .

Looking in the x-y plane at time-slices  $t_1, t_2, t_3, t_4$



The dephasing is reversed by the  $\pi_x$  pulse.

After the second time  $T$  the spins rephase and an echo signal will be seen.

